



THE UNIVERSITY  
of NORTH CAROLINA  
at CHAPEL HILL

# Time-Domain Astronomy with *SOFIA*: Results from Current Observations with FORCAST & Prospects with the Proposed New Instrument S3



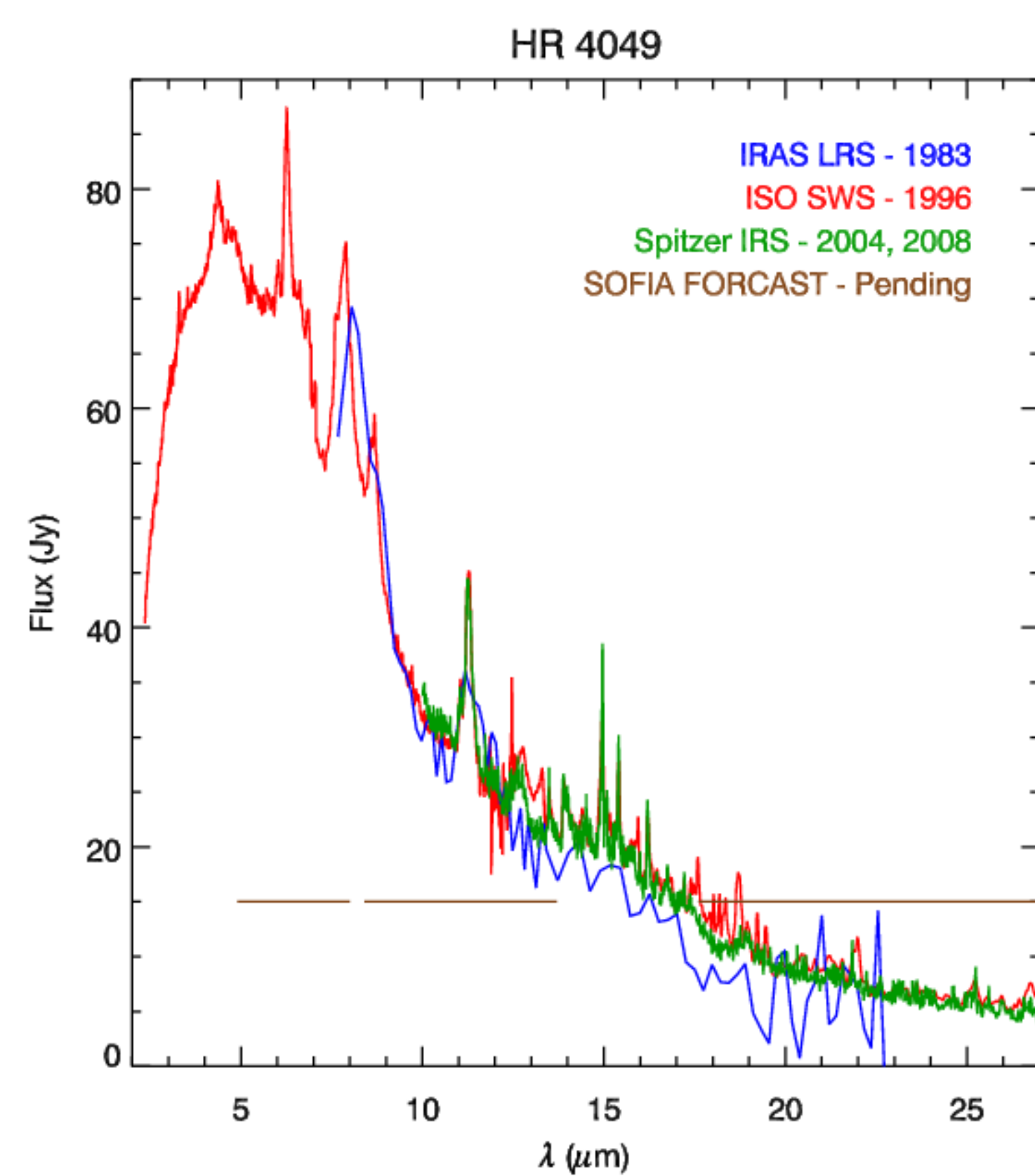
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The most recent Decadal Review identified "time-domain astronomy" as an important frontier for astrophysical investigations. A key advantage to *SOFIA* in this realm is its 20+ year lifetime which allows long-term scientific studies that are not possible with space-based infrared missions alone. Here, we present two time-domain projects enabled by *SOFIA*'s unique capabilities, one with current instrumentation - FORCAST, and the second with a proposed new instrument - S3.

## ‘Real-Time’ Evolution of Carbon-Rich Post-Asymptotic Giant Branch Stars 30+ Years: *IRAS* → *ISO* → *Spitzer* → *SOFIA*

**Project 1:** Obtain 5–37  $\mu\text{m}$  spectra with *SOFIA*'s FORCAST of carbon-rich post-Asymptotic Giant Branch (post-AGB), one of the most fleeting stages in a star's life.

- Most phases of stellar evolution occur on timescales of millions or billions of years, but the post-AGB phase, when a star rapidly transitions from a cool, dust-enshrouded object to a bare white dwarf illuminating a planetary nebula, **lasts only ~1,000-10,000 years**.
- We will compare the new FORCAST spectra to mid-IR spectra obtained at ~10–15 year intervals over the past 35 years with *Spitzer*, *ISO*, and *IRAS*.

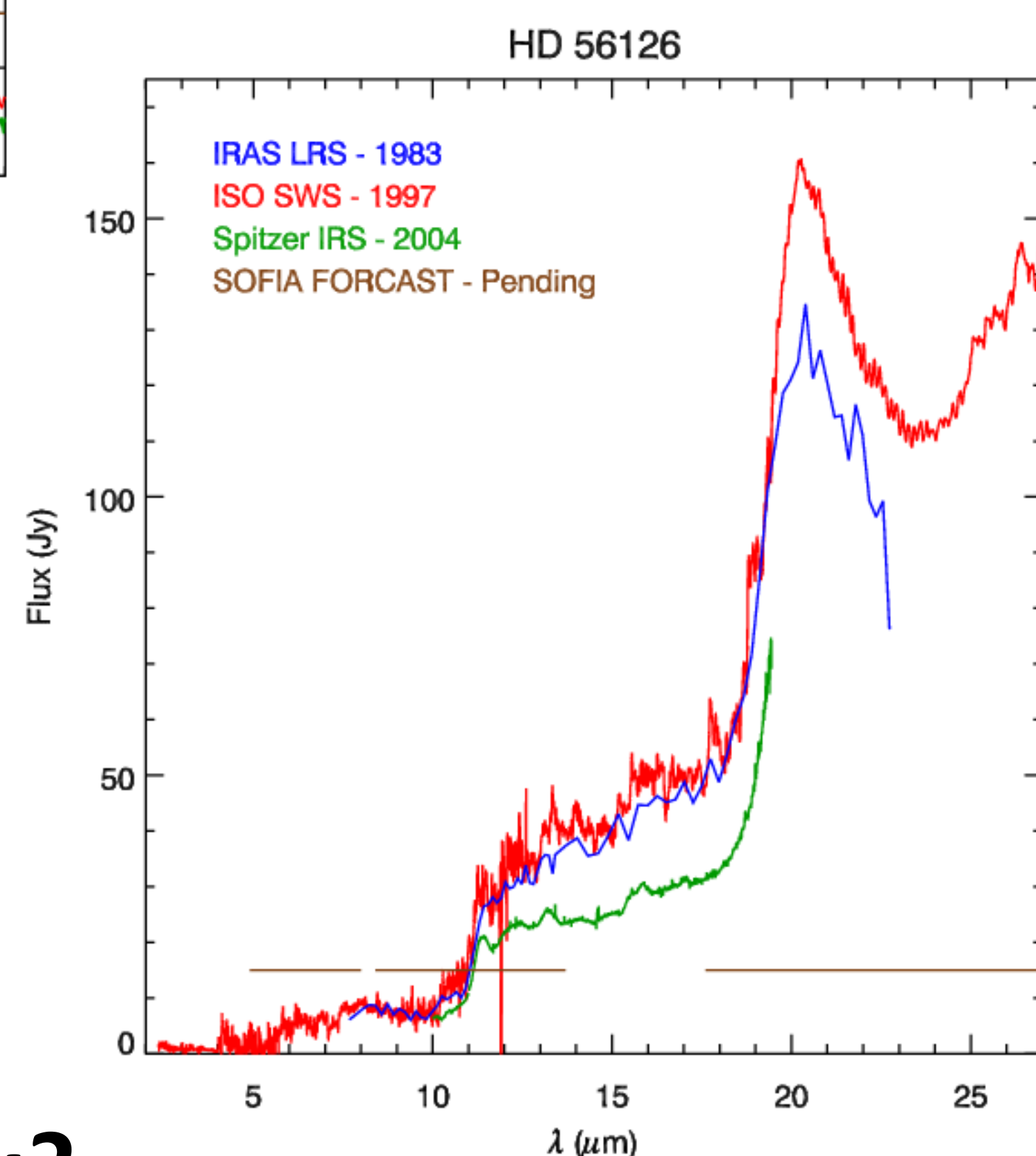


### ◀ Example 1: HR 4049

- Good agreement in shape & flux - even small bumps & wiggles for *ISO* & *Spitzer*
- Is the *IRAS/ISO* difference at ~7-8  $\mu\text{m}$  real?

### Example 2: HD 56126 ▶

- Intriguing differences in the 11-19  $\mu\text{m}$  *Spitzer* data
- Are the changes in the enigmatic 21  $\mu\text{m}$  feature real?



➤ **What will the FORCAST data show?**

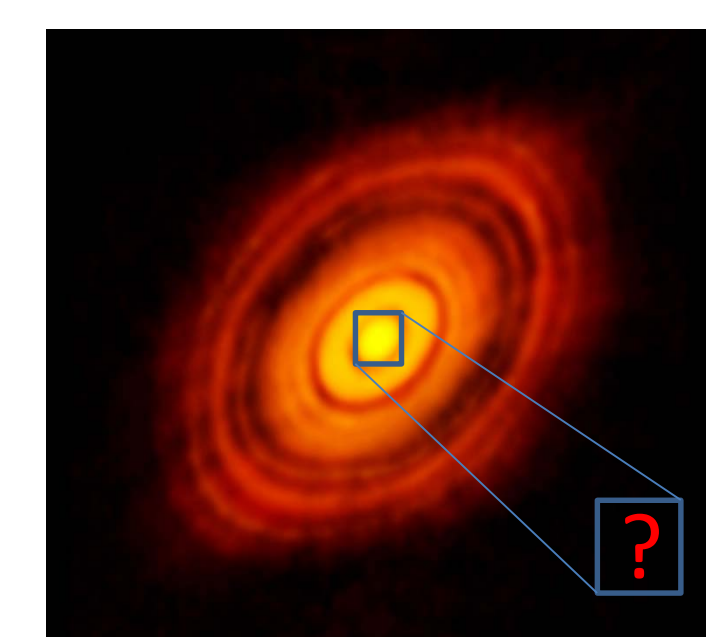
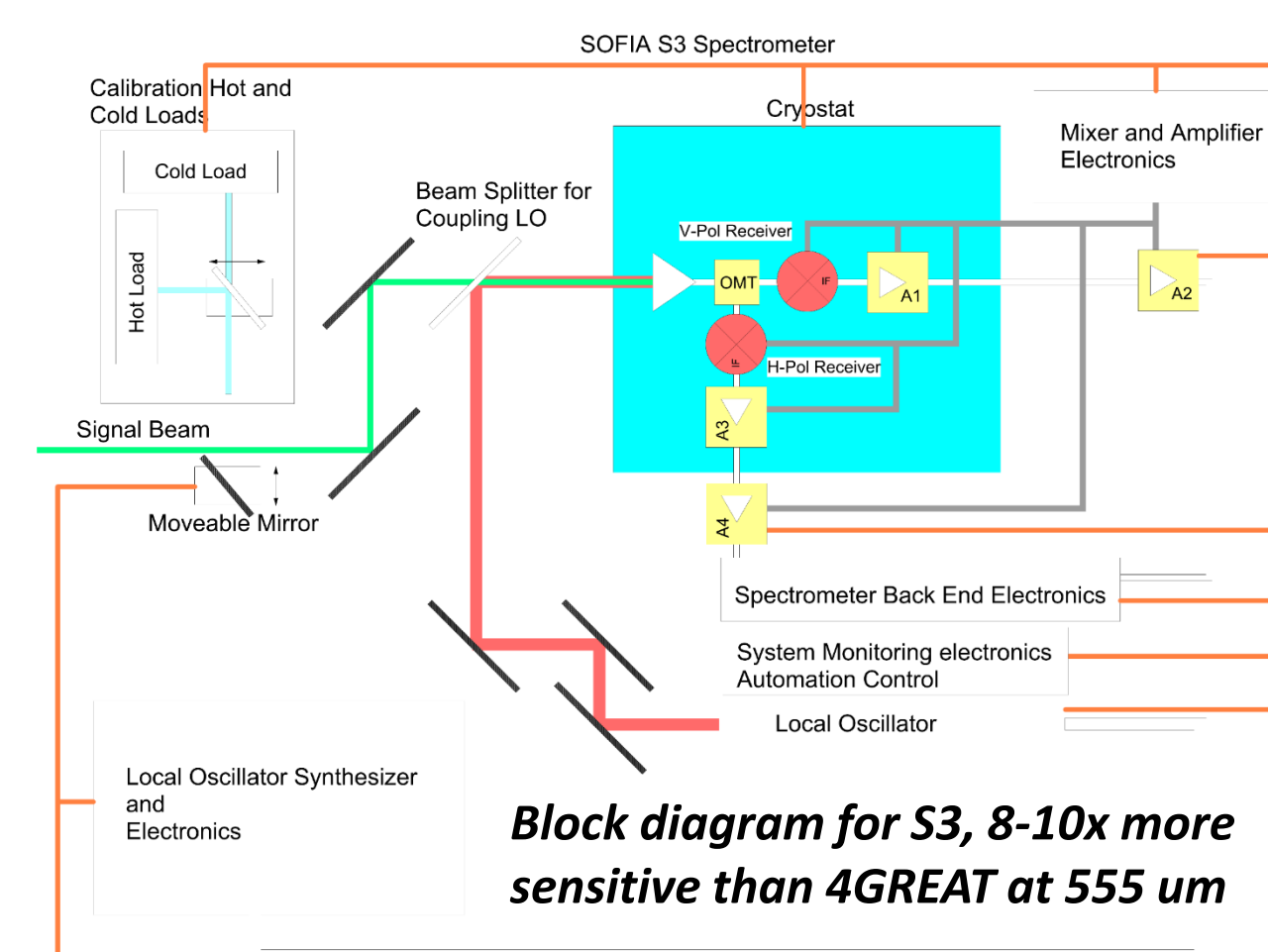
## Post-AGB Targets & Data

Star	Prior Datasets	SOFIA FORCAST	
		Observations	Data Delivery
RAFGL 618	<i>IRAS</i> LRS, <i>ISO</i> SWS	Aug. 22, 2018	Pending
IRAS 17441-2411	LRS, SWS,	Aug. 24, 2018	Pending
IRAS 19454+2920	LRS, SWS, <i>Spitzer</i> IRS	Aug. 25, 2018	Pending
IRAS 20000+3239	LRS, SWS, IRS	Aug. 25, 2018	Pending
V5112 Sgr	LRS, SWS	Aug. 29, 2018	Pending
HD 235858	LRS, SWS	Aug. 29, 2018	Pending
IRAS 19480+2504	LRS, SWS, IRS	Aug. 30, 2018	Pending
HD 44179	LRS, SWS	Pending	TBD
<b>HR 4049</b>	LRS, SWS, IRS	Pending	TBD
<b>HD 56126</b>	LRS, SWS, IRS	Pending	TBD

## Kinematic Imaging of Biogenic Molecules in Protoplanetary Disks with *S3: The Submillimeter Spectrometer for SOFIA*

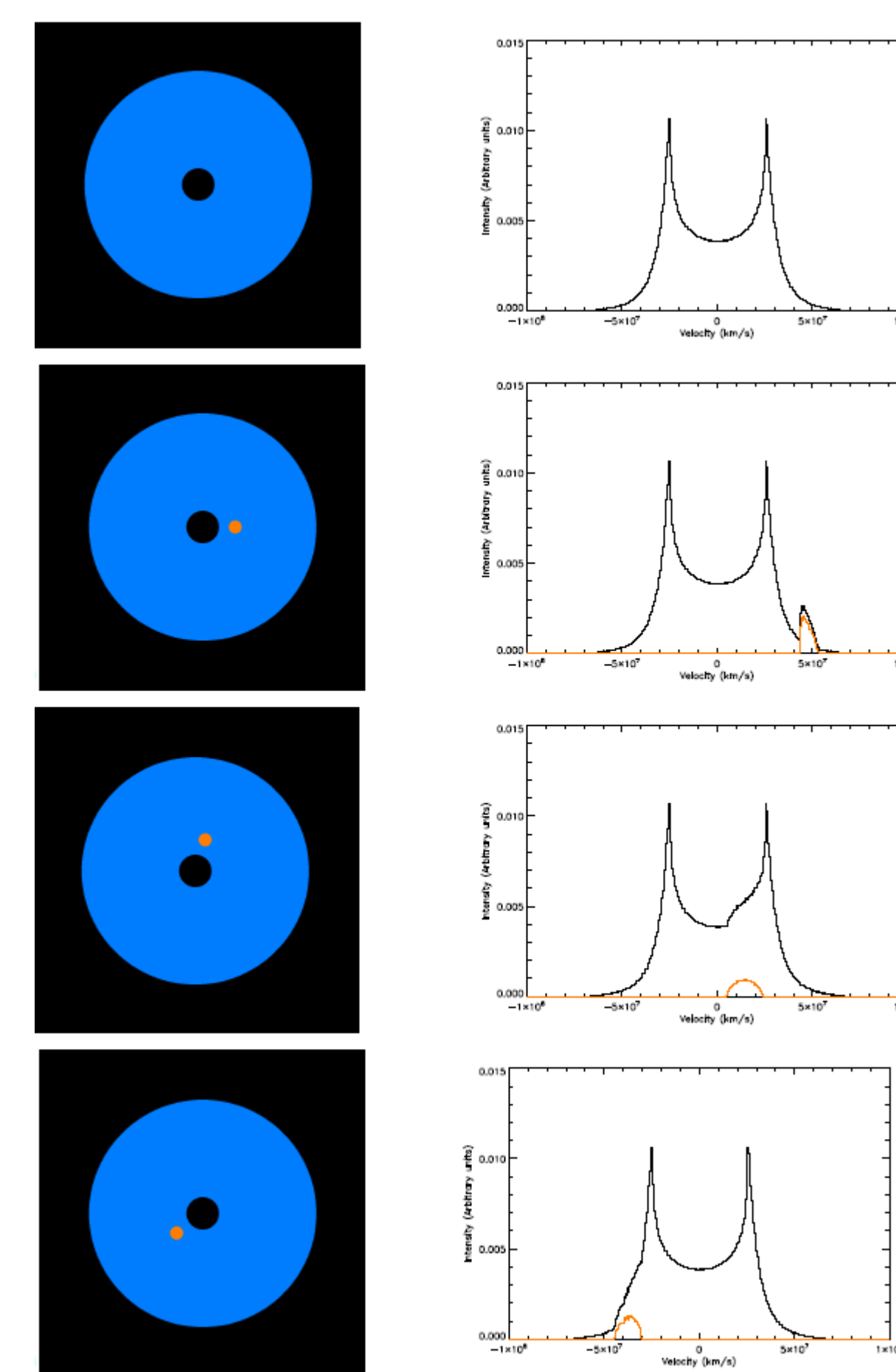
**Project 2:** Observe protoplanetary disks with a proposed new instrument: the high sensitivity, high spectral-resolution, heterodyne *Submillimeter Spectrometer for SOFIA*, S3

- Enabled by improved digital backend spectrometers, low noise amplifiers, local oscillators, and closed cycle refrigerators
- ❖ *Repeatedly* observe the line profiles of  $\text{NH}_3$  and isotopic  $\text{H}_2\text{O}$  in protoplanetary disks with S3 and their changes over the course of 3 years.
- ❖ Trace the structure and composition of the disk, constraining disk dynamics and planet formation models for each system.
- These observations by S3, **8-10x more sensitive than 4GREAT**, would allow Doppler tomography and similar techniques to kinematically image, for the first time, the **inner** regions of the planet-forming systems.



**S3: 2 unique new bands and a 3<sup>rd</sup> band an order of magnitude more sensitive than 4GREAT's**

### Kinematic Imaging



### ACKNOWLEDGMENTS and AFFILIATIONS

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